# Study of transport phenomena in plasma by extended irreversible thermodynamics 

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#### Abstract

One of most intellectually interesting problems in plasma physics is the problem of turbulence and the associated transport of the plasma properties including density, temperature and momentum. The aim of this paper is to determine the transport coefficients in plasma particularly electrical and thermal conductivities by Extended Irreversible Thermodynamics (EIT). Transport coefficients are determined for different types of particles electrons and ions.


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## 1. Introduction

Understanding turbulent transport in magnetized plasma is a subject of utmost importance for comprehending and optimizing experiments in present fusion devices and for designing future reactors, too [1-2].

The fundamental properties of plasma are markedly dependent upon the interactions of the plasma particles with the force fields existing inside it. Processes related to the transport of mass, momentum, energy and charges in plasma are generally called transport phenomena. There exist general equations describing these different phenomena, and the special effects are characterized by coefficients generally called transport coefficients [3].

Understanding and controlling the rate at which particles and heat escape from the reactor chamber is critical to the successful design and operation of a magnetic fusion device. In the early days of the fusion program, estimates of particle and heat transport based on simple collisional diffusion were made. However, these estimates were found to drastically under-predict the observed in experiments and the large measured transport was labeled "anomalous". Understanding this anomalous transport has been a primary goal of the fusion program ever since [4-5].

Recent descriptions of heat and particles transport in plasma have opened a promising field of application for Extended Irreversible Thermodynamics.

It has been shown that the so-called extended irreversible thermodynamics (EIT) attempts to cover those nonequilibrium situations which are supposed as not covered by local equilibrium assumption [6-8].

The basic features of this formalism and several applications are reviewed. Extended irreversible thermodynamics includes dissipative fluxes (heat flux, viscous pressure tensor, electric current) in the set of basic independent variables of the entropy [9].

This is achieved in EIT by enlarging the space of fundamental independent variables, such as the dissipative
fluxes; say the heat flux density, dissipative stress tensor etc [10-11]

A new formulation of nonequilibrium thermodynamics is based on the postulate that the entropy density ( S ) is a function of both ordinary thermodynamic variables and certain additional variables (heat flux, particle flux, etc.) [12-14]

The purpose of this paper is to determine the parallel and the perpendicular transport coefficients in plasma. In the second section we write the fundamental hypotheses of EIT and the corresponding evolution equations for the fluxes. In the third section, we develop the transport equations of particle and heat fluxes and we determine the parallel and perpendicular transport coefficients for multispecies of plasma (electrons and ions).

In the last section we comment the result obtained from graphical representation.

## 2. Generalized Gibbs equation

As in classical irreversible thermodynamics (CIT), the entropy and the Gibbs equation play a central role in extended irreversible thermodynamics (EIT). Here, it is assumed that the entropy will not only depend on the classical variable, namely the specific internal energy u, but in addition on the dissipative flux, so the generalized Gibbs equation takes the form [15-17]

$$
\begin{align*}
\mathrm{d} S^{\mathrm{a}}= & \frac{\mathrm{d} \mathrm{U}^{\mathrm{a}}}{\mathrm{~T}^{\mathrm{a}}}-\frac{\mu_{e}^{a}}{\mathrm{~T}^{\mathrm{a}}} \mathrm{~d} \mathrm{e}_{\mathrm{a}}-\frac{1}{\rho_{\mathrm{a}} \mathrm{~T}^{\mathrm{a}}}  \tag{1}\\
& {\left[\left(\alpha_{11}^{\mathrm{a}} \mathbf{q}^{\mathrm{a}}+\alpha_{12}^{a} \mathbf{j}^{\mathrm{a}}\right) \cdot \mathrm{dq}^{\mathrm{a}}+\left(\alpha_{21}^{a} \mathbf{q}^{\mathrm{a}}+\alpha_{22}^{\mathrm{a}} \mathbf{j}^{\mathrm{a}}\right) \mathrm{d} \mathbf{j}^{\mathrm{a}}\right] }
\end{align*}
$$

Where $U$ is the internal energy, $\mu_{\mathrm{e}}$ the chemical potential, $\alpha_{i j}$ phenomenological coefficient, $\mathrm{e}_{\mathrm{a}}$ total electric charge contributed by particle $\mathrm{a}, \mathrm{j}^{\mathrm{a}}$ particle flux, $\mathrm{q}^{\mathrm{a}}$ heat flux and $T_{a}$ is the absolute temperature of particle $\alpha$.
The balance equations of total electric charge contributed by particle a and internal energy are given by:

$$
\begin{equation*}
\rho^{\mathrm{a}} q_{\mathrm{t}} \mathrm{e}_{\mathrm{a}}=-\nabla \cdot \mathbf{j}^{\mathrm{a}} \tag{2}
\end{equation*}
$$

With $\rho^{a}$ is the total mass density of particle a

$$
\begin{equation*}
\rho^{\mathrm{a}} \partial_{\mathrm{t}} \mathrm{U}^{\mathrm{a}}=-\nabla \cdot \mathbf{q}^{\mathrm{a}}+\mathbf{j}^{\mathrm{a}} \cdot \mathrm{E} \tag{3}
\end{equation*}
$$

Here $E$ and $j^{a}$. $E$ are respectively the electric field and the joule heating term.
In virtue of the balance equations (2) and (3) for $U$ and $e^{a}$ , one obtains for the entropy balance

$$
\begin{align*}
& \rho^{\mathrm{a}} \partial_{\mathrm{t}} \mathrm{~S}+\nabla \cdot\left(\frac{1}{\mathrm{~T}_{\mathrm{a}}} \mathbf{q}^{\mathrm{a}}-\frac{\mu_{\mathrm{e}}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \mathbf{j}^{\mathbf{a}^{a}}\right)=\mathbf{q}^{\mathrm{a}} \cdot\left(\nabla \mathrm{~T}_{\mathrm{a}}^{-1}-\frac{\alpha_{11}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{q}^{\mathrm{a}}}{\mathrm{dt}}-\frac{\alpha_{21}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{j}^{\mathrm{a}}}{\mathrm{dt}}\right)+  \tag{4}\\
& \mathbf{j}^{\mathrm{a}} \cdot\left(\frac{\mathbf{E}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}}-\nabla \frac{\mu_{\mathrm{e}}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}}-\frac{\alpha_{12}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{q}^{\mathrm{a}}}{d \mathrm{t}}-\frac{\alpha_{22}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{j}^{\mathrm{a}}}{\mathrm{dt}}\right)
\end{align*}
$$

This equation can be cast in the general form of a balance equation

$$
\begin{equation*}
\rho \mathrm{S}=-\nabla \cdot \mathbf{j}^{\mathrm{a}}+\sigma^{\mathrm{a}} \tag{5}
\end{equation*}
$$

Where the quantity $\sigma^{\mathrm{a}}$ is the entropy production $\left(\sigma^{\mathrm{a}}>0\right)$ and $\mathrm{j}^{\mathrm{a}}$ is the entropy flux.
The entropy production obeys

$$
\begin{equation*}
\sigma^{\mathrm{a}}=\mu_{11} \mathbf{q}^{\mathrm{a}} \cdot \mathbf{q}^{\mathrm{a}}+\mu_{12} \mathbf{j}^{\mathrm{a}} \cdot \mathbf{q}^{\mathrm{a}}+\mu_{21} \mathbf{q}^{\mathrm{a}} \cdot \mathbf{j}^{\mathrm{a}}+\mu_{22} \mathbf{j}^{\mathrm{a}} \cdot \mathbf{j}^{\mathrm{a}} \tag{6}
\end{equation*}
$$

The coefficient $\mu \mathrm{ij}>0$ is a consequence of the entropy production with $\sigma^{\mathrm{a}}$ is positive $\left(\sigma^{\mathrm{a}}>0\right)$.

Considering equations (4) and (5), we obtain the simplest evolution equations for $\mathbf{q}^{\mathbf{a}}$ and $\mathbf{j}^{\mathrm{a}}$ compatible with a definite positive entropy production, one assumes linear relations between the thermodynamic forces and the fluxes $\mathbf{q}^{\mathrm{a}}$ and $\mathbf{j}^{\mathrm{a}}$. This result in

$$
\begin{align*}
& \frac{\mathbf{E}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}}-\nabla \frac{\mu_{\mathrm{a}}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}}-\frac{\alpha_{12}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{q}^{\mathrm{a}}}{\mathrm{dt}}-\alpha_{22}^{\mathrm{a}} \frac{\mathrm{~d} \mathbf{j}^{\mathrm{a}}}{\mathrm{dt}}=\mu_{21}^{\mathrm{a}} \mathbf{q}^{\mathrm{a}}+\mu_{22}^{\mathrm{a}} \mathbf{j}^{\mathrm{a}} \\
& \nabla \mathrm{~T}_{\mathrm{a}}^{-1}-\frac{\alpha_{11}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{q}^{\mathrm{a}}}{\mathrm{dt}}-\frac{\alpha_{21}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{j}^{\mathrm{a}}}{\mathrm{dt}}=\mu_{11}^{\mathrm{a}} \mathbf{q}^{\mathrm{a}}+\mu_{12}^{\mathrm{a}} \mathbf{j}^{\mathrm{a}} \tag{8}
\end{align*}
$$

Let us assume that $\nabla \mathrm{T}_{\mathrm{a}}^{-1}$ and $\frac{\mathrm{E}^{\mathrm{a}}}{\mathrm{T}_{\mathrm{a}}}-\nabla \frac{\mu_{\mathrm{a}}^{\mathrm{a}}}{\mathrm{T}_{\mathrm{a}}}$ vanish in system of equation (7) and (8), so that they refer to fluctuations near an equilibrium state. The equations (7) and (8) become:

$$
\begin{align*}
- & \frac{\alpha_{12}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{q}^{\mathrm{a}}}{\mathrm{dt}}-\frac{\alpha_{22}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{j}^{\mathrm{a}}}{\mathrm{dt}}=\mu_{21}^{\mathrm{a}} \mathbf{q}^{\mathrm{a}}+\mu_{22}^{\mathrm{a}} \mathbf{j}^{\mathrm{a}}  \tag{9}\\
& -\frac{\alpha_{11}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{q}^{\mathrm{a}}}{\mathrm{dt}}-\frac{\alpha_{21}^{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}} \frac{\mathrm{~d} \mathbf{j}^{\mathrm{a}}}{\mathrm{dt}}=\mu_{11}^{\mathrm{a}} \mathbf{q}^{\mathrm{a}}+\mu_{12}^{\mathrm{a}} \mathbf{j}^{\mathrm{a}} \tag{10}
\end{align*}
$$

After resolution of equations (9) and (10) we have

$$
\begin{array}{ll}
\lambda^{\mathrm{b}}=\frac{5 \mathrm{n}_{\mathrm{b}} \mathrm{k}_{\mathrm{b}}^{2} \mathrm{~T}_{\mathrm{b}}}{2 \mathrm{~m}_{\mathrm{b}}} \tau^{\mathrm{b}} & , \sigma^{\mathrm{b}}=\frac{\mathrm{n}_{\mathrm{b}} \mathrm{e}_{\mathrm{b}}^{2} \mathrm{c}^{\mathrm{b}}}{\mathrm{~m}_{\mathrm{b}}} \\
\Pi^{\mathrm{b}}=\frac{\pi^{2} \mathrm{k}_{\mathrm{B}} \mathrm{~T}_{\mathrm{b}}^{2}}{2 \mathrm{e}_{\mathrm{b}} \varepsilon_{\mathrm{F}}^{\mathrm{b}}} & , \varepsilon^{\mathrm{b}}=\frac{\pi^{2} \tau_{\mathrm{b}}^{2} \mathrm{k}_{\mathrm{b}}^{2}}{2 \mathrm{e}_{\mathrm{b}}^{2} \varepsilon_{\mathrm{F}}^{\mathrm{b}}} \\
\mu^{\mathrm{b}}=\mathrm{k}_{\mathrm{B}} \mathrm{~T}_{\mathrm{b}} \log \left(\frac{1}{\mathrm{n}_{\mathrm{b}}}\left(\frac{2 \pi \mathrm{k}_{\mathrm{B}} \mathrm{~T}_{\mathrm{b}}}{\mathrm{~m}_{\mathrm{b}}}\right)^{\frac{3}{2}}\right.
\end{array}
$$

Where $\tau_{b} \quad n_{b}$ and $k_{B}$ are respectively the relaxation time of particle $b$, the particle $b$ density and the Boltzmann constant.
Determination of the coefficients $\alpha_{\mathrm{ij}}^{\mathrm{ab}}$ [18]

$$
\begin{array}{ll}
\alpha_{11}^{\mathrm{ab}}=\frac{\mathrm{k}_{\mathrm{B}}}{\left\langle\delta \mathbf{q}^{\mathrm{a}} \delta \mathbf{q}^{\mathrm{b}}\right\rangle} & ; \alpha_{12}^{\mathrm{ab}}=\frac{\mathrm{k}_{\mathrm{B}}}{\left\langle\delta \mathbf{q}^{\mathrm{a}} \delta \mathbf{j}^{\mathrm{b}}\right\rangle}  \tag{17}\\
\alpha_{21}^{\mathrm{ab}}=\frac{\mathrm{k}_{\mathrm{B}}}{\left\langle\delta \mathbf{j}^{\mathrm{a}} \delta \mathbf{q}^{\mathrm{b}}\right\rangle} \quad ; \quad \alpha_{22}^{\mathrm{ab}}=\frac{\mathrm{k}_{\mathrm{B}}}{\left\langle\delta \mathbf{j}^{\mathrm{a}} \delta \mathbf{j}^{\mathrm{b}}\right\rangle}
\end{array}
$$

With $\partial \mathbf{q}^{a}$ is the fluctuation of the heat flux, $\partial \mathbf{j}^{\mathrm{a}}$ is the fluctuation of the particle flux.

Determination of $\alpha_{i j}^{a b}$ :
The fluctuations of the heat flux are given

$$
\begin{equation*}
\delta \mathbf{q}^{\mathrm{a}}=\int\left(\frac{1}{2} \mathrm{mC}^{2}-\frac{5}{2} \mathrm{k}_{\mathrm{B}} \mathrm{~T}\right) \mathrm{C} \delta \mathrm{f}^{\mathrm{a}} \tag{18}
\end{equation*}
$$

Using this expression, frequently called the subtracted heat flux, to compute the second moments of the fluctuations, one finds that:

$$
\begin{gather*}
\left\langle\delta \mathbf{q}^{\mathrm{a}} \delta \mathbf{q}^{\mathrm{b}}\right\rangle=\int \mathrm{dc} \int \mathrm{dc}\left(\frac{1}{2} \mathrm{~m}_{\mathrm{a}} \mathrm{C}^{2}-\frac{5}{2} \mathrm{k}_{\mathrm{B}} \mathrm{~T}_{\mathrm{a}}\right) \mathrm{C}\left(\frac{1}{2} \mathrm{~m}_{\mathrm{b}} \mathrm{C}^{\prime 2}-\frac{5}{2} \mathrm{k}_{\mathrm{B}} \mathrm{~T}_{\mathrm{b}}\right) \mathrm{C}^{\prime}  \tag{19}\\
\left\langle\delta \mathrm{f}^{\mathrm{a}}(\mathrm{C}) \delta \mathrm{f}^{\mathrm{b}}\left(\mathrm{C}^{\prime}\right)\right\rangle \\
\left\langle\delta \mathbf{j}^{\mathrm{a}} \delta \mathbf{q}^{\mathrm{b}}\right\rangle=\int \mathrm{dc} \int \mathrm{~d} \mathrm{c}^{\prime}\left(\frac{1}{2} \mathrm{C}\left(\frac{1}{2} \mathrm{~m}_{\mathrm{b}} \mathrm{C}^{\prime 2}-\frac{5}{2} \mathrm{k}_{\mathrm{B}} \mathrm{~T}_{\mathrm{b}}\right) \mathrm{C} \mathrm{C}\right.  \tag{20}\\
\\
\left\langle\delta \mathrm{f}^{\mathrm{a}}(\mathrm{C}) \delta \mathrm{f}^{\mathrm{b}}\left(\mathrm{C}^{\prime}\right)\right\rangle  \tag{21}\\
\left\langle\delta \mathbf{q}^{\mathrm{a}} \delta \mathbf{j}^{\mathrm{b}}\right\rangle=\mathrm{e}_{\mathrm{b}} \int \mathrm{dc} \int \mathrm{dc}\left(\frac{1}{2} \mathrm{C}\left(\frac{1}{2} \mathrm{~m}_{\mathrm{b}} \mathrm{C}^{\prime 2} \frac{5}{2} \mathrm{~K}_{\mathrm{b}} \mathrm{~T}_{\mathrm{b}}\right) \mathrm{CC}^{\prime}\right) \\
\left\langle\delta \mathrm{f}^{\mathrm{a}}(\mathrm{C}) \delta \mathrm{f}^{\mathrm{b}}\left(\mathrm{C}^{\prime}\right)\right\rangle \tag{22}
\end{gather*}
$$

Where $\mathrm{C}=\mathrm{c}-\mathrm{v}$ the particle velocity relative to the mean motion, with $c$ is the velocity of a particle, $v$ is the mean velocity.

$$
\begin{equation*}
\left\langle\delta \mathrm{f}^{\mathrm{a}}(\mathrm{C}) \delta \mathrm{f}^{\mathrm{b}}\left(\mathrm{C}^{\prime}\right)\right\rangle=\frac{1}{\mathrm{v}} \mathrm{f}_{\mathrm{eq}}(\mathrm{C}) \delta(\mathrm{C}-\mathrm{C}) \tag{23}
\end{equation*}
$$

Where V is the volume of the system

The local equilibrium Maxwell-Boltzmann distribution function $\mathrm{f}_{\mathrm{eq}}$

$$
\mathrm{f}_{\mathrm{eq}}=\mathrm{n}\left(\frac{\mathrm{~m}}{2 \pi \mathrm{~K}_{\mathrm{b}} \mathrm{~T}}\right)^{3 / 2} \exp \left(-\frac{\mathrm{mC}^{2}}{2 \mathrm{~K}_{\mathrm{B}} \mathrm{~T}}\right)
$$

After calculations, we found that


$$
\alpha_{22}^{a b}=\frac{12 V_{B}}{\left(\mathrm{e}_{\mathrm{a}} \mathrm{n}_{\mathrm{b}} \mathrm{n}_{\mathrm{a}} \mathrm{~T}_{\mathrm{b}} \sqrt{\pi} \beta_{\mathrm{a}}^{3}\right)}\left(\frac{2 \pi^{2} \mathrm{~T}_{\mathrm{b}}}{\mathrm{~m}_{\mathrm{b}}}\right)^{\frac{3}{2}}\left(1+\left(\frac{\beta_{\mathrm{a}}}{\beta_{\mathrm{b}}}\right)^{2}\right)^{\frac{3}{2}}
$$

With: $\quad \beta_{\mathrm{a}}=\left(\frac{\mathrm{m}_{\mathrm{a}}}{2 \mathrm{~T}_{\mathrm{a}}}\right)^{1 / 2} \quad$ and $\quad \beta_{\mathrm{b}}=\left(\frac{\mathrm{m}_{\mathrm{b}}}{2 \mathrm{~T}_{\mathrm{b}}}\right)^{1 / 2}$
Here $\mathrm{m}_{\mathrm{a}}, \mathrm{n}_{\mathrm{a}}, \mathrm{V}$ and $\mathrm{a}_{\mathrm{a}^{\prime}}$ are respectively the mass, the density of particle a, volume of the system and the charge of particle a .

The state of the plasma after a short transition time remains close to the local plasma equilibrium. For this reason, the local plasma equilibrium will be a reference state. The distribution function can the conveniently be written in the form: [17]

$$
\begin{equation*}
\mathrm{f}^{\mathrm{a}}=\mathrm{f}_{0}^{\mathrm{a}}+\mathrm{f}_{1}^{\mathrm{a}} \tag{24}
\end{equation*}
$$

With $f_{1}^{a}$ is a deviation of distribution function, and then can be expanded in a series of irreducible Hermite polynomial as

$$
\begin{equation*}
\mathrm{f}_{1}^{\mathrm{a}}=\sum_{\mathrm{n}=0} \mathrm{~h}^{\mathrm{a}(2 \mathrm{n}+1)} \mathrm{H}_{\mathrm{r}}^{(2 \mathrm{n}+1)}(\mathrm{v}) \mathrm{f}_{0}^{\mathrm{a}}(\mathrm{v}) \tag{25}
\end{equation*}
$$

With $h^{\mathrm{a}(2 \mathrm{n}+1)}$ the hermitien moment and $\mathrm{H}_{\mathrm{r}}^{(2 \mathrm{n}+1)}$ the irreducible Hermite polynomials
We limited in the 13 Moment approximation ( $\mathrm{n}=1$ ) so

$$
\begin{equation*}
\mathrm{f}_{1}^{\mathrm{a}}=\left(\mathrm{h}_{\mathrm{r}}^{\mathrm{a}(\mathrm{l})} \mathrm{H}_{\mathrm{r}}^{\mathrm{a}(1)}+\mathrm{h}_{\mathrm{r}}^{\mathrm{a}(3)} \mathrm{H}_{\mathrm{r}}^{\mathrm{a}(3)}\right) \mathrm{f}_{0}^{\mathrm{a}} \tag{26}
\end{equation*}
$$

With

$$
\begin{align*}
& \mathrm{H}_{\mathrm{r}}^{\mathrm{a}(1)}=\sqrt{2} \beta_{\mathrm{a}} \mathrm{C}  \tag{36}\\
& \mathrm{H}_{\mathrm{r}}^{\mathrm{a}(\mathrm{a})}=\sqrt{\frac{1}{5}} \beta_{\mathrm{a}} \mathrm{C}\left[2\left(\beta_{\mathrm{a}} \mathrm{C}\right)^{2}-5\right]
\end{align*}
$$

$$
\begin{equation*}
\mathrm{G}_{\mathrm{r}}^{\mathrm{a}(\mathrm{a})}=-\tau_{\mathrm{a}} \sqrt{\frac{5}{2}}\left(\frac{\mathrm{~m}_{\mathrm{a}}}{\mathrm{~T}_{\mathrm{a}}}\right)^{1 / 2} \frac{1}{\mathrm{~m}_{\mathrm{a}}} \nabla_{\mathrm{r}}\left(\mathrm{~T}_{\mathrm{a}}\right) \tag{34}
\end{equation*}
$$

Considering equations (12) and (13) the evolution equations of particle and heat flux has the form

$$
\begin{array}{r}
\partial_{\mathrm{t}} \mathbf{j}^{\mathrm{a}}=\sum_{\mathrm{b}=\mathrm{e}, \mathrm{i}}\left(\mathrm{k}_{1}^{\mathrm{ab}} \cdot \mathbf{q}^{\mathrm{b}}+\mathrm{k}_{2}^{\mathrm{ab}} \mathbf{j}^{\mathrm{b}}\right)  \tag{35}\\
\partial_{\mathrm{t}} \mathbf{q}^{\mathrm{a}}=\sum_{\mathrm{b}=\mathrm{e}, \mathrm{i}}\left(\mathrm{k}_{1}^{\mathrm{ab}} \cdot \mathbf{q}^{\mathrm{b}}+\mathrm{k}_{4}^{\mathrm{ab}} \mathbf{j}^{\mathrm{b}}\right)
\end{array}
$$

So the dimensionless equations of particle and heat fluxes in parallel direction become:

$$
\begin{gather*}
\partial_{\mathrm{t}} \mathrm{~h}_{/ /}^{\mathrm{b}(1)}=\sum_{\mathrm{b}=\mathrm{e}, \mathrm{i}}\left(\mathrm{k}_{1}^{\mathrm{ab}} \mathrm{~h}_{/ /}^{\mathrm{b}(3)}+\mathrm{k}_{4}^{\mathrm{ab}} \mathrm{~h}_{/ /}^{\mathrm{b}(1)}\right)  \tag{37}\\
\partial_{\mathrm{t}} \mathrm{~h}_{/ /}^{\mathrm{a}(3)}=\sum_{\mathrm{b}=\mathrm{e}, \mathrm{i}}\left(\mathrm{k}_{1}^{\mathrm{ab}} \mathrm{~h}_{/ /}^{\mathrm{b}(3)}+\mathrm{k}_{4}^{\mathrm{ab}} \mathrm{~h}_{/ /}^{\mathrm{b}(1)}\right) \tag{38}
\end{gather*}
$$

With:

$$
\begin{aligned}
& \mathrm{k}_{1}^{\mathrm{ab}}=\sqrt{\frac{5}{2}} \frac{\mathrm{n}_{\mathrm{b}}}{\mathrm{n}_{\mathrm{a}}} \mathrm{~T}_{\mathrm{b}}\left(\frac{\mathrm{~m}_{\mathrm{a}} \mathrm{~T}_{\mathrm{b}}}{\mathrm{~m}_{\mathrm{b}} \mathrm{~T}_{\mathrm{a}}}\right)^{\frac{1}{2}} \mathrm{k}_{1}^{\mathrm{ab}} \\
& \mathrm{k}_{2}^{\mathrm{ab}}=\frac{\mathrm{n}_{\mathrm{b}}}{\mathrm{n}_{\mathrm{a}}}\left(\frac{\mathrm{~m}_{\mathrm{a}} \mathrm{~T}_{\mathrm{b}}}{\mathrm{~m}_{\mathrm{b}} \mathrm{~T}_{\mathrm{a}}}\right)^{\frac{1}{2}} \mathrm{k}_{2}^{\mathrm{ab}} \\
& \mathrm{k}_{3}^{\mathrm{ab}}=\sqrt{\frac{5}{2}} \frac{\mathrm{n}_{\mathrm{b}}}{\mathrm{n}_{\mathrm{a}}} \mathrm{~T}_{\mathrm{b}}\left(\frac{\mathrm{~m}_{\mathrm{a}} \mathrm{~T}_{\mathrm{b}}}{\mathrm{~m}_{\mathrm{b}} \mathrm{~T}_{\mathrm{a}}}\right)^{\frac{1}{2}} \mathrm{k}_{3}^{\mathrm{ab}} \\
& \mathrm{k}_{4}^{\mathrm{ab}}=\frac{\mathrm{n}_{\mathrm{b}}}{\mathrm{n}_{\mathrm{a}}}\left(\frac{\mathrm{~m}_{\mathrm{a}} \mathrm{~T}_{\mathrm{b}}}{\mathrm{~m}_{\mathrm{b}} \mathrm{~T}_{\mathrm{a}}}\right)^{\frac{1}{2}} \mathrm{k}_{4}^{\mathrm{ab}}
\end{aligned}
$$

The dimensionless evolution equation (31) of particle flux in parallel direction became

$$
\begin{equation*}
\tau^{\mathrm{a}} \partial_{\mathrm{t}} \mathrm{~h}_{/ /}^{\mathrm{a}(1)}+\mathrm{h}_{/ /}^{\mathrm{a}(1)}=\mathrm{G}_{/ /}^{\mathrm{a}(1)} \tag{39}
\end{equation*}
$$

Where the term of magnetic field vanish
For electrons (e), the equation (39) become

$$
\begin{equation*}
\tau^{\mathrm{e}} \partial_{\mathrm{t}} \mathrm{~h}_{/ /}^{\mathrm{e}(1)}+\mathrm{h}_{/ /}^{\mathrm{e}(1)}=\mathrm{G}_{/ /}^{\mathrm{e}(1)} \tag{40}
\end{equation*}
$$

The expression of electron relaxation time $\tau^{e}$ is [19]

$$
\begin{equation*}
\tau^{e}=\frac{3 m_{e}^{1 / 2} T_{e}^{3 / 2}}{4 \sqrt{2 \pi} Z^{2} e^{4} n_{i} \ln \Lambda} \tag{41}
\end{equation*}
$$

Where $\ln \Lambda=\ln \frac{3 / 2\left(T_{e}+T_{i}\right) \lambda_{D}}{Z e^{2}}$ the coulomb logarithm, and $\lambda_{D}$ is the Debye length.

We replace $\quad \tau^{e} \partial_{t} h_{r}^{e(1)}$ in equation (40) by its expression from equation (37), we find:

$$
\begin{equation*}
\tau^{e} \sum_{b=e, i}\left(k_{1}^{\prime a b} h_{/ /}^{b(3)}+k_{2}^{\prime a b} h_{/ /}^{b(1)}\right)+h_{/ /}^{e(1)}=G_{/ /}^{e(1)} \tag{42}
\end{equation*}
$$

This is equivalent to:

$$
\begin{align*}
& \tau^{e} k_{1}^{\text {'ee }} h_{\|}^{e(3)}+\tau^{e} k_{2}^{\dot{e e}} h_{/ /}^{e(1)}+\tau^{e} k_{1}^{\dot{e i}} h_{/ /}^{i(3)}+\tau^{e} k_{2}^{\text {ei }} h_{/ /}^{i(1)}+h_{/ /}^{e(1)}  \tag{43}\\
& \quad=G_{/ l}^{e(1)}
\end{align*}
$$

Considering Fourier transformed

$$
\begin{aligned}
& \partial_{t} h_{/ /}^{b(1)}=i \omega h_{/ /}^{b(1)} \\
& \partial_{t} h_{/ /}^{b(3)}=i \omega h_{/ /}^{b(3)}
\end{aligned}
$$

We place in the asymptotic limit where we neglect $\partial_{t} h_{r}$, with we take $\omega=0$, the evolutions equations (43) reduces to

$$
\begin{equation*}
h_{/ /}^{e(1)}=L_{11}^{e e} G_{/ /}^{e(1)}+L_{13}^{e e} G_{/ /}^{e(3)}+L_{11}^{e i} G_{/ /}^{i(1)}+L_{13}^{e i} G_{/ /}^{i(3)} \tag{44}
\end{equation*}
$$

Similar for ions (i) we determine $h_{/ /}^{i(1)}$

$$
\begin{equation*}
h_{/ /}^{i(1)}=L_{11}^{i e} G_{/ /}^{e(1)}+L_{13}^{i e} G_{/ /}^{e(3)}+L_{11}^{i i} G_{/ /}^{i(1)}+L_{13}^{i i} G_{/ /}^{i(3)} \tag{45}
\end{equation*}
$$

The $L_{i j}^{a a}$ coefficients with ( $\mathrm{a}=\mathrm{e}, \mathrm{i}$ and $\mathrm{i}=1 \mathrm{j}=1,3$ ) are the pure transport coefficients and $L_{i j}^{a b}(a, b=e, i)$ with ( $\mathrm{a} \neq$ b)) the mixed transport coefficients.

Similar of particle flux we develop now the dimensionless evolution equation of heat flux of multispecies of plasma so we have
$h_{/ /}^{e(3)}=L_{31}^{e e} G_{/ /}^{e(1)}+L_{33}^{e e} G_{/ /}^{e(3)}+L_{31}^{e i} G_{/ /}^{i(1)}+L_{33}^{e i} G_{/ /}^{i(3)}$
$h_{/ /}^{i(3)}=L_{31}^{i e} G_{/ /}^{e(1)}+L_{33}^{i e} G_{/ /}^{e(3)}+L_{31}^{i i} G_{/ /}^{i(1)}+L_{33}^{i i} G_{/ /}^{i(3)}$

## Identification of the transport coefficients

The transport matrix $L_{i j}^{a b}$ (whose coefficients are the transport coefficients) has the characteristic Onsager symmetry, which reduces here to the simple matrix symmetry $L_{i j}^{a b}=L_{j i}^{a b}$ so

## The pure transport coefficients have the form

$$
\begin{aligned}
L_{11}^{e e} & =\frac{1}{1+\tau^{e} k_{2}^{\prime e e}} & , \quad L_{13}^{e e}=-\frac{\tau^{e} k_{1}^{\prime e e}}{1+\tau^{e} k_{2}^{\prime e e}} \\
L_{33}^{i i} & =-\frac{1}{1+\tau^{i} k_{2}^{\prime i i}} \quad, & L_{13}^{i i}=-\frac{\tau^{i} k_{1}^{\prime i i}}{1+\tau^{i} k_{2}^{i i i}}, \\
L_{33}^{e e} & =-\frac{1}{1+\tau^{e} k_{2}^{\prime e e}} &
\end{aligned}
$$

With the expression of ion relaxation time is

$$
\begin{equation*}
\tau^{i}=\frac{3 m_{i}^{1 / 2} T_{i}^{3 / 2}}{4 \sqrt{2 \pi} Z^{2} e^{4} n_{i} \ln \Lambda} \tag{48}
\end{equation*}
$$

Where $\mathrm{m}_{\mathrm{i}}, \mathbf{Z}_{i}$ and $\mathrm{T}_{\mathrm{i}}$ are respectively the mass of ion, the charge and the ion temperature.

## The mixed transport coefficients

$$
\begin{aligned}
L_{11}^{e i} & =-\frac{\tau^{e} k_{2}^{\prime e i}}{1+\tau^{e} k_{2}^{\prime e e}}, \quad L_{31}^{e i}=-\frac{\tau^{e} k_{4}^{\prime e i}}{1+\tau^{e} k_{3}^{\prime e e}} \\
L_{33}^{e i} & =-\frac{\tau^{e} k_{3}^{\prime e i}}{1+\tau^{e} k_{3}^{\prime e e}}
\end{aligned}
$$

## 4. The perpendicular transport coefficients of plasma

In this paragraph we study the transport phenomena in presence of a constant magnetic field (B) where the particles (electrons (e), ions (i)), move perpendicular to the magnetic fields.
$\mathbf{B}$ is parallel to the Z axis so we project the dimensionless evolution equations in the x and y direction

## For particle fluxes

We determine now the dimensionless evolution equation of particle flux of multispecies of plasma so we have

$$
\begin{align*}
h_{x}^{e(1)}= & L_{11}^{e e} G_{x}^{e(1)}+L_{13}^{e e} G_{x}^{e(3)}+L_{11}^{e i} G_{x}^{i(1)}+L_{13}^{e i} G_{x}^{i(3)} \\
& +\left(\frac{-e B \tau_{e}}{m_{e} c}\right) h_{y}^{e(1)}  \tag{49}\\
h_{y}^{e(1)}= & L_{11}^{e e} G_{y}^{e(1)}+L_{13}^{e e} G_{y}^{e(3)}+L_{11}^{e i} G_{y}^{i(1)}+L_{13}^{e i} G_{y}^{i(3)} \\
& +\left(\frac{-e B \tau_{e}}{m_{e} c}\right) h_{x}^{e(1)} \tag{50}
\end{align*}
$$

We suppose
$x_{a}=\frac{e_{a} B \tau_{a}}{m_{a} c}, \quad x_{a}=\Omega_{a} \tau_{a} \quad$ With $\Omega_{a}=\frac{e_{a} B}{m_{a} c}$
is the Larmor frequency of species $a, a=(e, i)$ and $B$ is the magnetic field
After introducing (50) in (49) we use
For electrons
$h_{x}^{e(1)}=\frac{L_{11}^{e e}}{1+x_{e}^{2}} G_{x}^{e(1)}+\frac{L_{13}^{e e}}{1+x_{e}^{2}} G_{x}^{e(3)}+\frac{L_{11}^{e i}}{1+x_{e}^{2}} G_{x}^{i(1)}$
$+\frac{L_{13}^{e i}}{1+x_{e}^{2}} G_{x}^{i(3)}+x_{e} \frac{L_{11}^{e e}}{1+x_{e}^{2}} G_{y}^{e(1)}+x_{e} \frac{L_{13}^{e e}}{1+x_{e}^{2}} G_{y}^{e(3)}+x_{e} \frac{L_{11}^{e i}}{1+x_{e}^{2}} G_{y}^{i(1)}+$
$x_{e} \frac{L_{13}^{e i}}{1+x_{e}^{2}} G_{y}^{i(3)}$

$$
\begin{align*}
& h_{y}^{e(1)}=\frac{L_{11}^{e e}}{1+x_{e}^{2}} G_{y}^{e(1)}+\frac{L_{13}^{e e}}{1+x_{e}^{2}} G_{y}^{e(3)}+\frac{L_{11}^{e i}}{1+x_{e}^{2}} G_{y}^{i(1)}+\frac{L_{13}^{e i}}{1+x_{e}^{2}} G_{y}^{i(3)} \\
& -x_{e} \frac{L_{11}^{e e}}{1+x_{e}^{2}} G_{x}^{e(1)}-x_{e} \frac{L_{13}^{e e}}{1+x_{e}^{2}} G_{x}^{e(3)}-x_{e} \frac{L_{11}^{e i}}{1+x_{e}^{2}} G_{x}^{i(1)}- \\
& \quad x_{e} \frac{L_{13}^{e i}}{1+x_{e}^{2}} G_{x}^{i(3)} \tag{52}
\end{align*}
$$

For ions

$$
\begin{align*}
& h_{x}^{i(1)}=\frac{L_{11}^{i e}}{1+x_{i}^{2}} G_{x}^{e(1)}+\frac{L_{13}^{i e}}{1+x_{i}^{2}} G_{x}^{e(3)}+\frac{L_{11}^{i i}}{1+x_{i}^{2}} G_{x}^{i(1)} \\
& +\frac{L_{13}^{i i}}{1+x_{i}^{2}} G_{x}^{i(3)}+x_{i} \frac{L_{11}^{i e}}{1+x_{i}^{2}} G_{y}^{e(1)}+x_{i} \frac{L_{13}^{i e}}{1+x_{i}^{2}} G_{y}^{e(3)}+x_{i} \frac{L_{11}^{i i}}{1+x_{i}^{2}} G_{y}^{i(1)}+ \\
& x_{i} \frac{L_{13}^{i i}}{1+x_{i}^{2}} G_{y}^{i(3)}  \tag{53}\\
& h_{y}^{i(1)}=\frac{L_{11}^{i e}}{1+x_{i}^{2}} G_{y}^{e(1)}+\frac{L_{13}^{i e}}{1+x_{i}^{2}} G_{y}^{e(3)}+\frac{L_{11}^{i i}}{1+x_{i}^{2}} G_{y}^{i(1)}+\frac{L_{13}^{i i}}{1+x_{i}^{2}} G_{y}^{i(3)} \\
& -x_{i} \frac{L_{11}^{i e}}{1+x_{i}^{2}} G_{x}^{e(1)}-x_{i} \frac{L_{13}^{i e}}{1+x_{i}^{2}} G_{x}^{e(3)}-x_{i} \frac{L_{11}^{i i}}{1+x_{i}^{2}} G_{x}^{i(1)}- \\
& \quad x_{i} \frac{L_{13}^{i i}}{1+x_{i}^{2}} G_{x}^{i(3)} \tag{54}
\end{align*}
$$

## For heat fluxes

Similar of particle flux we develop now the dimensionless evolution equation of heat flux of multispecies of plasma so we have

$$
\begin{align*}
h_{x}^{e(3)}= & L_{31}^{e e} G_{x}^{e(1)}+L_{33}^{e e} G_{x}^{e(3)}+L_{31}^{e i} G_{x}^{i(1)}+L_{33}^{e i} G_{x}^{i(3)} \\
& +\left(\frac{-e B \tau_{e}}{m_{e} c}\right) h_{y}^{e(3)}  \tag{55}\\
h_{y}^{e(3)}= & L_{31}^{e e} G_{y}^{e(1)}+L_{33}^{e e} G_{y}^{e(3)}+L_{31}^{e i} G_{y}^{i(1)}+L_{33}^{e i} G_{y}^{i(3)} \\
& +\left(\frac{-e B \tau_{e}}{m_{e} c}\right) h_{x}^{e(3)} \tag{56}
\end{align*}
$$

After introducing (56) in (55) we use
For electrons

$$
\begin{align*}
& h_{x}^{e(3)}=\frac{L_{31}^{e e}}{1+x_{e}^{2}} G_{x}^{e(1)}+\frac{L_{33}^{e e}}{1+x_{e}^{2}} G_{x}^{e(3)}+\frac{L_{31}^{e i}}{1+x_{e}^{2}} G_{x}^{i(1)} \\
& +\frac{L_{33}^{e i}}{1+x_{e}^{2}} G_{x}^{i(3)}+x_{e} \frac{L_{31}^{e e}}{1+x_{e}^{2}} G_{y}^{e(1)}+x_{e} \frac{L_{33}^{e e}}{1+x_{e}^{2}} G_{y}^{e(3)}+x_{e} \frac{L_{31}^{e i}}{1+x_{e}^{2}} G_{y}^{i(1)}+ \\
& x_{e} \frac{L_{33}^{e i}}{1+x_{e}^{2}} G_{y}^{i(3)}  \tag{57}\\
& h_{y}^{e(3)}=\frac{L_{31}^{e e}}{1+x_{e}^{2}} G_{y}^{e(1)}+\frac{L_{33}^{e e}}{1+x_{e}^{2}} G_{y}^{e(3)}+\frac{L_{31}^{e i}}{1+x_{e}^{2}} G_{y}^{i(1)}+\frac{L_{33}^{e i}}{1+x_{e}^{2}} G_{y}^{i(3)} \\
& -x_{e} \frac{L_{31}^{e e}}{1+x_{e}^{2}} G_{x}^{e(1)}-x_{e} \frac{L_{33}^{e e}}{1+x_{e}^{2}} G_{x}^{e(3)}-x_{e} \frac{L_{31}^{e i}}{1+x_{e}^{2}} G_{x}^{i(1)}- \\
& \quad x_{e} \frac{L_{33}^{e i}}{1+x_{e}^{2}} G_{x}^{i(3)} \tag{58}
\end{align*}
$$

For ions

$$
\begin{align*}
& h_{x}^{i(3)}=\frac{L_{31}^{i e}}{1+x_{i}^{2}} G_{x}^{e(1)}+\frac{L_{33}^{i e}}{1+x_{i}^{2}} G_{x}^{e(3)}+\frac{L_{31}^{i i}}{1+x_{i}^{2}} G_{x}^{i(1)} \\
& +\frac{L_{33}^{i i}}{1+x_{i}^{2}} G_{x}^{i(3)}+x_{i} \frac{L_{31}^{i e}}{1+x_{i}^{2}} G_{y}^{e(1)}+x_{i} \frac{L_{33}^{i e}}{1+x_{i}^{2}} G_{y}^{e(3)}+x_{i} \frac{L_{31}^{i i}}{1+x_{i}^{2}} G_{y}^{i(1)}+ \\
& x_{i} \frac{L_{33}^{i j}}{1+x_{i}^{2}} G_{y}^{i(3)}  \tag{59}\\
& h_{y}^{i(3)}=\frac{L_{31}^{i e}}{1+x_{i}^{2}} G_{y}^{e(1)}+\frac{L_{33}^{i e}}{1+x_{i}^{2}} G_{y}^{e(3)}+\frac{L_{31}^{i i}}{1+x_{i}^{2}} G_{y}^{i(1)}+\frac{L_{33}^{i i}}{1+x_{i}^{2}} G_{y}^{i(3)} \\
& -x_{i} \frac{L_{31}^{i e}}{1+x_{i}^{2}} G_{x}^{e(1)}-x_{i} \frac{L_{33}^{i e}}{1+x_{i}^{2}} G_{x}^{e(3)}-x_{i} \frac{L_{31}^{i i}}{1+x_{i}^{2}} G_{x}^{i(1)}- \\
& x_{i} \frac{L_{33}^{i i}}{1+x_{i}^{2}} G_{x}^{i(3)} \tag{60}
\end{align*}
$$

## 5. Result and discussion

In Figs. 1-4 we plotted the perpendicular electrical conductivity, the thermoelectric conductivity, and electron and ion thermal conductivities as function of $\mathrm{x}_{\mathrm{a}}(\mathrm{a}=\mathrm{e}, 1)$.


Fig. 1. Perpendicular electrical conductivity as function of $\mathrm{K}_{\mathbf{z}}$.

This plot show that the electrical conductivity decreases with increasing of $\mathrm{x}_{\mathrm{E}}$


Fig. 2. Perpendicular electron thermal conductivity as function of $\mathrm{X}_{\mathrm{E}}$.

We notice that the perpendicular electron thermal conductivity decreases with increasing of $\mathrm{x}_{\mathbf{8}}$.


Fig. 3. Perpendicular thermoelectric conductivity as function of $\mathrm{X}_{\mathrm{E}}$.

We notice that the perpendicular thermoelectric conductivity increases with increasing of $\mathrm{x}_{\mathrm{e}}$.


Fig. 4. Perpendicular ions thermal conductivity as function of $\mathrm{x}_{\mathrm{i}}$.

In this curve we plotted the perpendicular ions thermal conductivity as function $\mathrm{x}_{\mathrm{i}}$ of. This plot shows that the perpendicular ions thermal conductivity decreases with increasing of $\mathrm{x}_{i}$

The perpendicular transport coefficients are monotonously decreasing function of $\mathrm{x}_{\mathrm{a}}$. So for a very strong magnetic field, the particles would stick to the field lines, and there would be no transport in any direction perpendicular to (B). This situation is opposed by the collisions the latter make the particles jump from one field line to another, thus making a perpendicular transport possible. In the perpendicular direction the collisions favor the transport thus for large values of the parameter
$x_{a}=\Omega_{a} \tau_{a} \quad$ i.e. for large magnetic field the coefficients are decreasing functions of $\mathrm{x}_{\mathrm{a}}$. This situation is clearly illustrated in Figs. 1-4.

## 6. Conclusion

In this paper we have been interested to calculate by Extended Irreversible Thermodynamics the transport coefficients like electrical and thermal conductivities after developing transport equations of dissipative fluxes like particles and heat fluxes of multispecies plasma (electron and ions).

In absence of magnetic field all the transport coefficient are proportional to the relaxation time this important characteristic can be easily understood. It implies that as the collision frequency increases the transport coefficients decrease in order words, the collisions tend to oppose the transport of matter and energy; they act as an obstacle to the free flow of these
quantities. This result is in perfect agreement with kinetic theory result [20-21].

The asymptotic perpendicular transport coefficients are proportional to the collision frequency this property vividly illustrates that the collisions oppose the parallel transport but favor the perpendicular one. It may be said that in plasma in presence of a constant magnetic field, when the collision frequency is increased

And the equality such coefficients, which has been obtained here from purely EIT, is supported by kinetic theory.

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## References

[1] W. Horton, Rev. Mod. Phy, 71, 735 (1999).
[2] H. Mordman, J. Weiland, Nuclear fusion 27941 (1987).
[3] X. Garbet, at all, Plasma Phys. Control Fusion, 46, 2004.
[4] J. B. Lister, F. Hofmann, J. M. Moret, Fusion Technology, 32, num. 3, (1997).
[5] W. Horton, R. D. Estes, Plasma Phys, 22, 663 (1980).
[6] A. Jarmen, P. Anderson, J. Weiland, Nuclear Fusion 27, 941 (1987).
[7] F. Miskane, A. Dezairi, X. Garbet, Phys. Plasma, 7, 4197 (2000).
[8] D. Jou, J. Casas-V'azquez, G. Lebon, Extended irreversible thermodynamics (Springer-Verlag, Berlin, 1993, 1996.
[9] R. E. Nettleton, S. L. Sobolev, J. Non-Equilib. Thermodyn. 20, 205, 297 (1995); 21, 1 (1996).
[10] P. Salamon, S. Sieniutycz (eds), Extended thermodynamic systems (Taylor and Francis, New York, 1992).
[11] D. Jou, J. Casas-Vàzquez, G. Lebon, Extended Irreversible Thermodynamics, second ed., Springer, Berlin, 1996.
[12] D. Jou, J. Casas-V'azquez, Phys. Rev. E45, 8371 (1992).
[13] D. Jou, J. Casas-V'azquez, G. Lebon, Rep. Prog. Phys. 51, 1105 (1988).
[14] L. S. Garcia-Colin, F. J. Uribe, J. Non Equilib. Thermodyn. 16, 89 (1991).
[15] R. E. Nettleton, Can. J. Phys. 72, 106 (1994).
[16] D. Jou, J. Casas-V'azquez, Phys. Rev. E48, 3201 (1993).
[17] T. Dedeurwaerdere, J. Casas. Vàzques, D. Jou, Phys. Rev. E 53 (1996).
[18] D. Jou, J. Casas-V'azquez, G. Lebon, Extended irreversible thermodynamics third, Revised and enlarged edition, Springer.
[19] R. Balescu, Equilibrium and nonequilibrium statistical Mechanics Wiley-Interscience, 1975.
[20] HIRSHMAN, S. P. Phys. Fluids 19, 155 (1976).
[21] R. Balescu, Phys Fluids B4, 91(1992).

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